



Alfvén Eigen Mode and Energetic Particle Studies in Alcator C-Mod

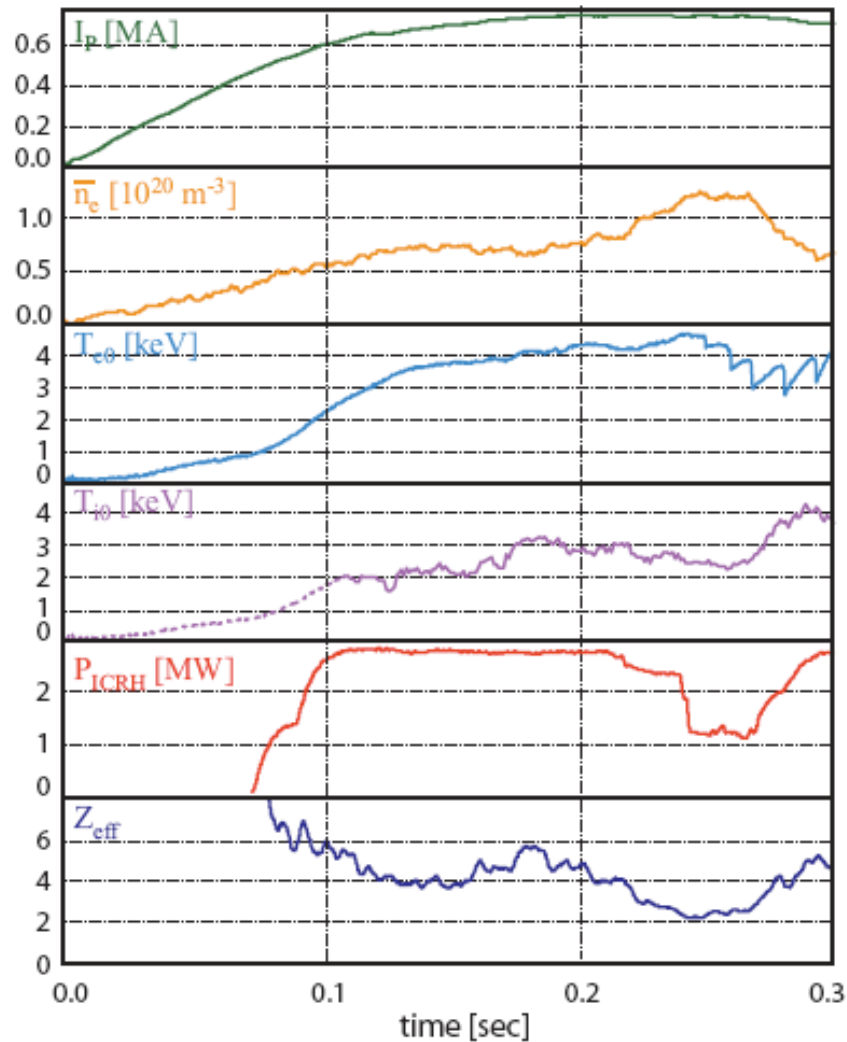
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In collaboration with E. Edlund, J. Dorris, R. Parker, C. Rost, P. Woskov, S. Wukitch
and the Alcator team

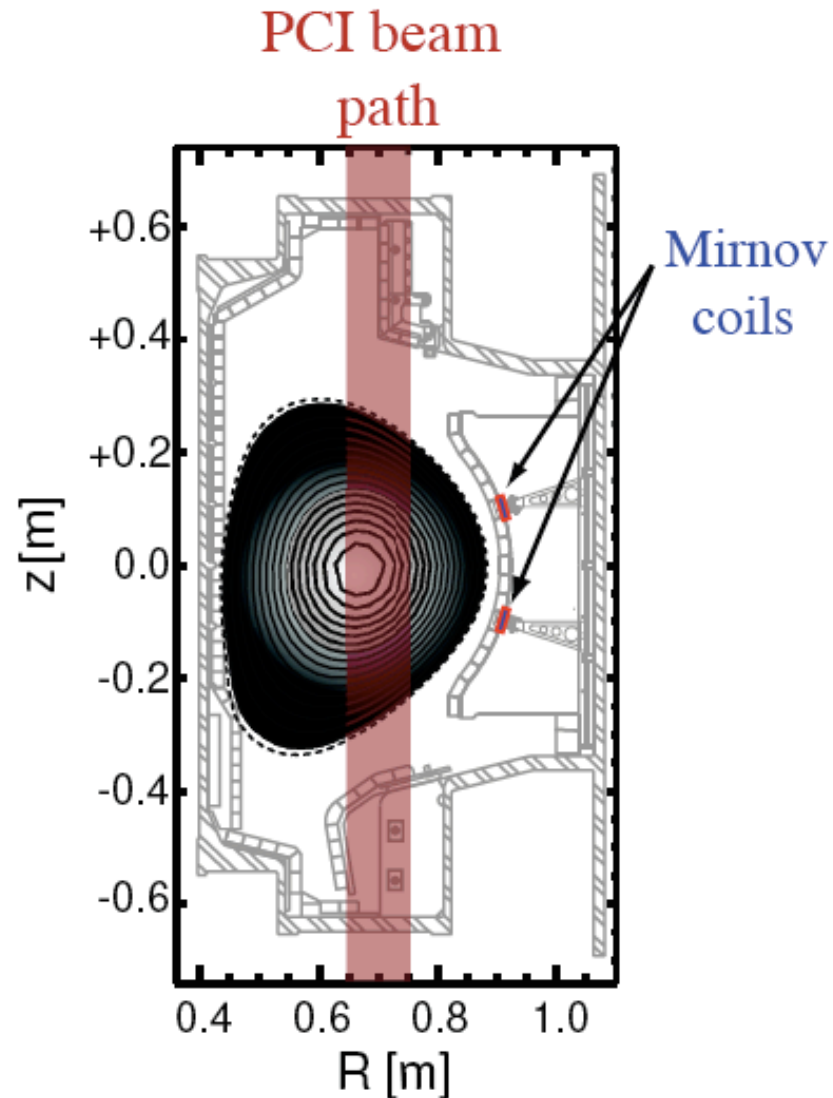
Recent Reference: E. Edlund, M. Porkolab, G.J. Kramer, et al,
revised MS resubmitted to PPCF, 7.2010.

Presented at GSEP Workshop, 8.10.2010, General Atomics

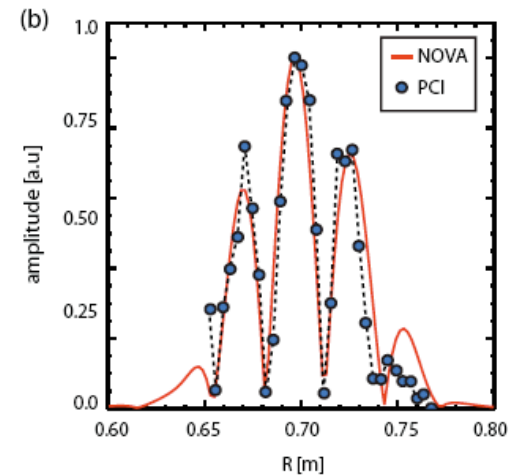
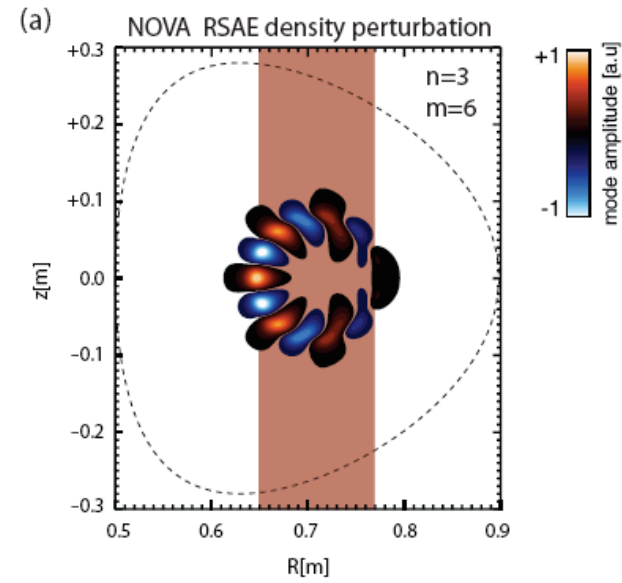
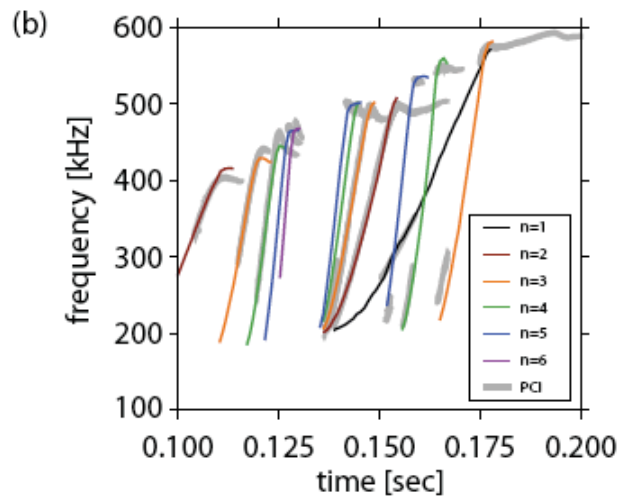
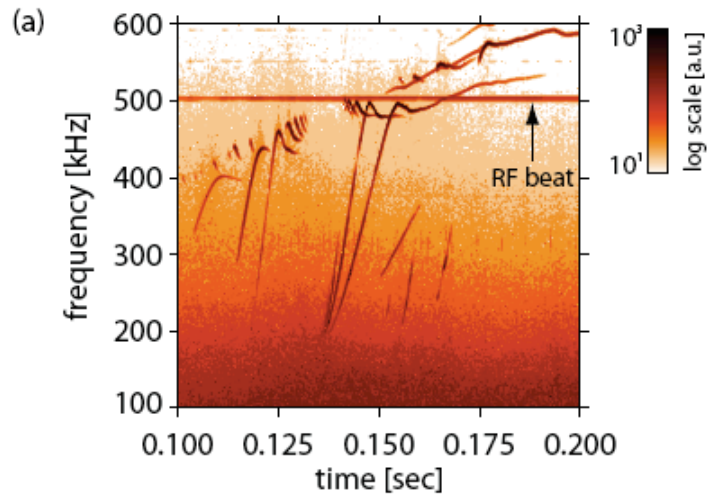
Typical time traces of current ramping studies for RSAE in C-Mod :
Reversed shear q-profile during t= 0.05 sec to 0.20 sec;
Also find RSAE in between sawtooth crashes: 0.3 sec + : RS q profile ?



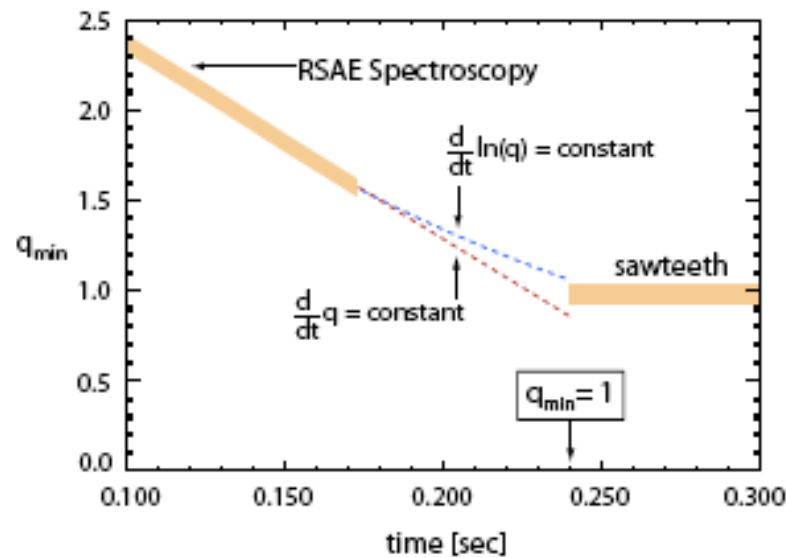
Phase Contrast Imaging (PCI) in Alcator C-Mod images RSAE in the plasma core in a more sensitive manner than the Mirnov coils in the plasma edge; synthetic diagnostic installed in NOVA



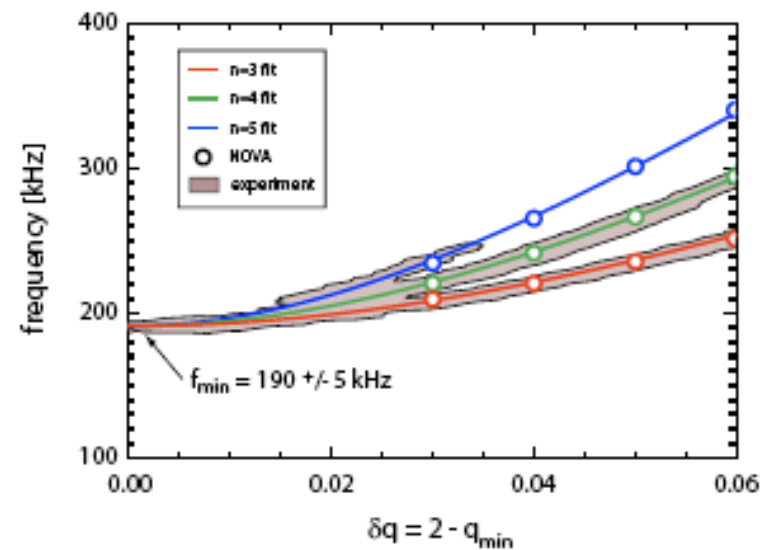
RSAE in the initial phase of the ramp-up allows deduction of $q_{\min}(t)$ and the reconstruction of the eigenmodes is in good agreement with PCI synthetic diagnostic in NOVA



Variation of the minimum RSAE frequency with q_{\min} and limiting values near $q=1$ between sawtooth crash phenomena

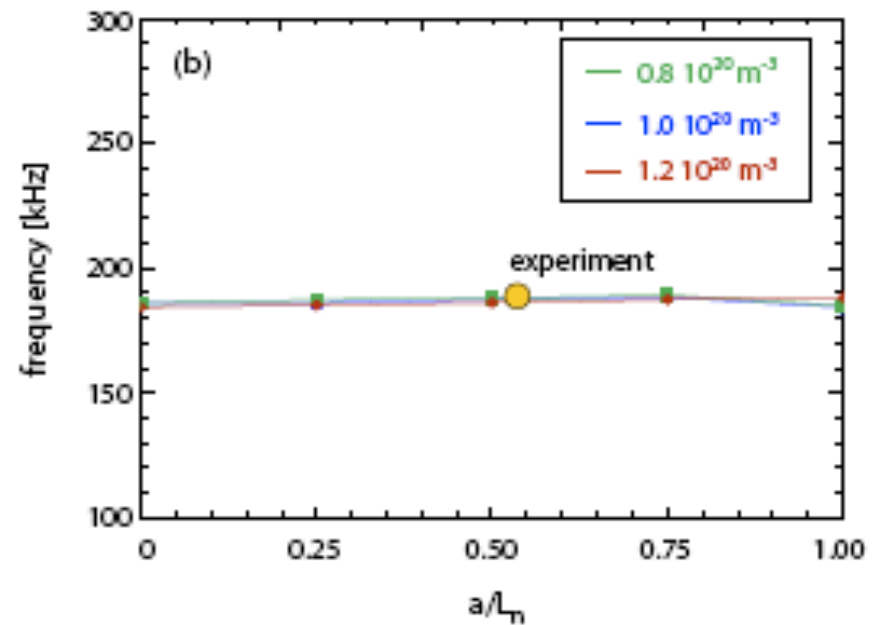
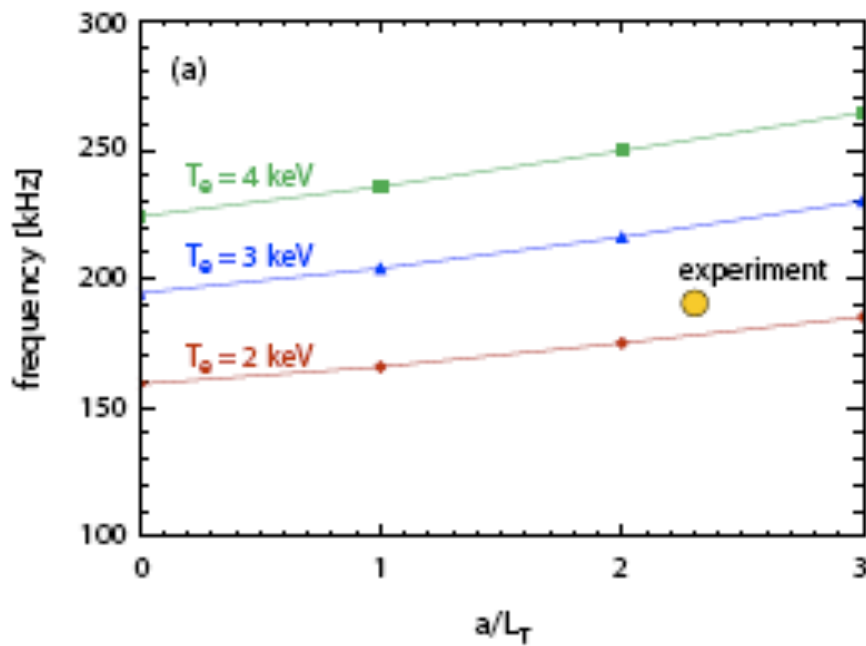


MHD spectroscopy during the current ramp and ultimately sawteeth provide limiting values of q_{\min}

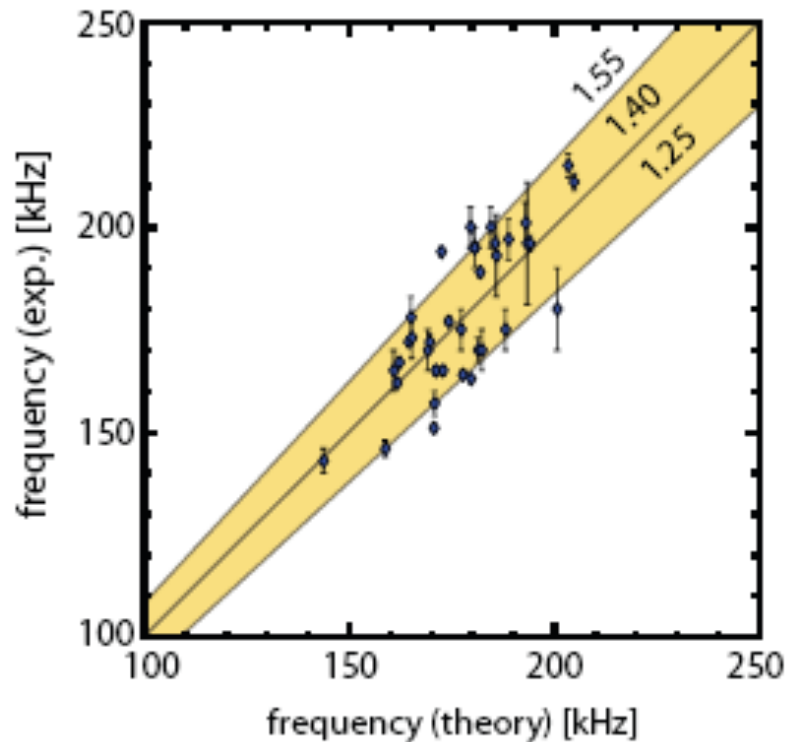


Nova modeling of the "trend" in the minimum frequencies

Parameter scans in NOVA show that the minimum RSAE frequency shows a strong T_e dependence, a weaker $\text{grad}(T_e)$ dependence, and a very weak n_e or $\text{grad}(n_e)$ dependence



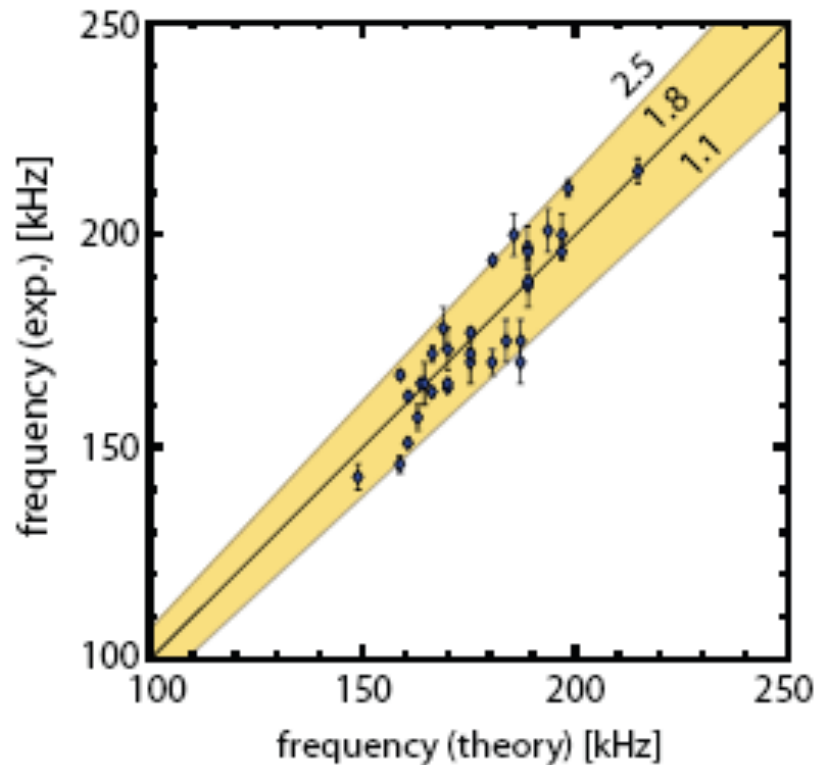
Measured minimum frequency data is compared with NOVA modeling to derive gamma. The best value is 1.40 ± 0.15 , close to $3/2$ predicted by C.F. McKee and E. Zweibel, *Astrophysical Journal*, 1995, and not $5/3$; (is there a theoretical basis for this gamma?)



$$\omega_{NOVA} = 100 T_e^{1/2} \sqrt{\gamma + 0.15} \frac{r}{L_T}$$

$$p = (\gamma - 1)U, \text{ dictates that } \gamma = 3/2$$

Minimum frequency data compared to formula of Winsor, Breizman, etc.
 The best fit appears to be gamma-ion about 1.8, which is very close to $7/4 = 1.75$ and within experimental error indistinguishable.



$$\omega_{min}^2 \approx \frac{2T_e}{m_i R_0^2} (1 + \gamma_i \tau)$$

To compare with NOVA modeling, based on MHD, we propose that

$$\gamma_{eff} = \frac{1 + \gamma_i \tau}{1 + \tau}$$

▪ **Future Plans include TAE Excitation in JET (8 antenna array, MIT (Woskov, Porkolab) , CRPP (Fasoli) , JET Collaboration in 2011; in 2011 MIT would buy 8 new independent amplifiers, 1.0 KW each, 10 kHz to 1 MHz bandwidth (OFES funded project);**

Due to delays in the JET program (rebuilding JET internals for ITER relevant tiles) we have decided to purchase 2 amplifiers now and start tuning and coupling experiments on C-Mod with existing double strap antennas (winter, 2010, 2011)